

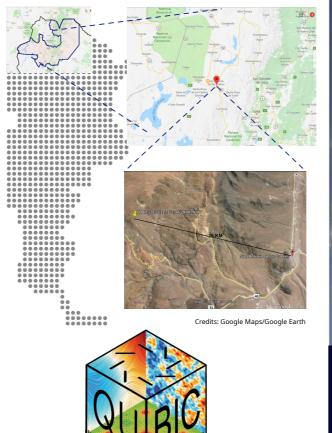
The first module of QUBIC is placed in Alto Chorillo (Salta, Argentina), 16 km from San Antonio de los Cobres, at ~5000 m.a.s.l.

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QUBIC, in its final version, will include 6 telescopes installed on an area of approximately 10,000 m². This region is destined to become an "Astronomical Pole" and it is part of the "Salta Ventana al Universo" program.

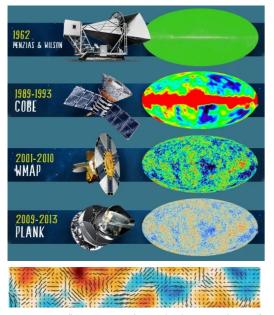


https://www.qubic.org.ar



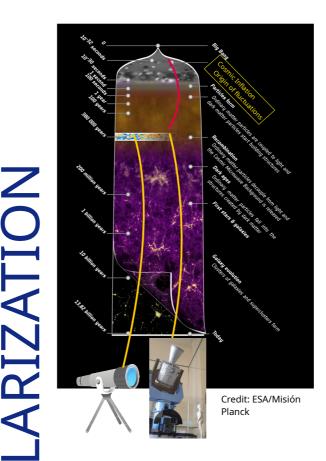


Q&U Bolometric Interferometer for Cosmology



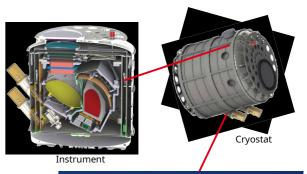
Temperature fluctuations (color) and polarization (lines) of the CMB on a small patch of sky Credit: ESA/ Planck Mission

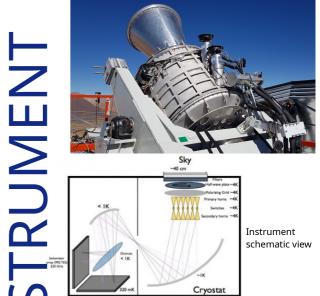
The cosmic microwave background (CMB) is the remaining radiation, relic of the origin of the Universe, which was left from the decoupling between the photons and the matter, an event that took place during the early Universe, 380,000 years after the Big Bang. Its detection in 1964 and its very fluctuations small temperature observed in 1992 thanks to the satellites COBE, both discoveries celebrated by Nobel prizes, are among the main scientific advances of the 20th century. The deepened studies in the 21st century by WMAP and Planck space telescopes, have led us to a deeper understanding of the history of our Universe.



 QUBIC is a cosmology experiment that aims to measure the polarization of cosmic microwave background radiation with great precision.

This study would offer us an invaluable insight into what happened immediately after the Big Bang; for example, it would allow us to test inflation theories that claim that there was an extremely rapid exponential expansion during the first 10⁻³⁸ seconds of the Universe, an extremely small fraction of time.





The very weak polarized signal of the CMB requires even more complex and sensitive experiments than those developed so far, in order to detect it. challenge, То face this the Collaboration in the QUBIC Project has developed completely а new instrument based on the groundbreaking concept of bolometric interferometry, which combines the high sensitivity of bolometric detectors with the purity of interferometry.